

## On Neutrino in Deformed Relativity

Lukasz Andrzej Glinka\*

Member of the American Association of International Researchers. New York, USA

### Editorial

Glinka [1] has considered the algebraic deformation of Special Relativity, which led to the equation

$$(\gamma^\mu \hat{p} \mu + mc^2 + \sqrt{\alpha} \frac{c}{\hbar} \ell \gamma^5 \hat{p}^2) \psi = 0, \quad (1)$$

Which can be easily transformed into the pair of Dirac equations

$$(\gamma^\mu \hat{p} \mu - M \pm c^2) \psi = 0, \quad (2)$$

Where  $M \pm$  are the emerging mass matrices, which can be written in the form

$$M \pm = \mu R \frac{1 + \gamma^5}{2} + \mu L \frac{1 - \gamma^5}{2}, \quad (3)$$

$$\mu R = -\frac{1}{c^2} \left( \frac{\varepsilon}{2} \pm \sqrt{\varepsilon^2 - 4\varepsilon mc^2 - 4E^2} \right), \quad (4)$$

$$\mu L = \frac{1}{c^2} \left( \frac{\varepsilon}{2} \pm \sqrt{\varepsilon^2 - 4\varepsilon mc^2 - 4E^2} \right), \quad (5)$$

Where  $\mu R, \mu L$  reals and the energy parameter are

$$\varepsilon = \frac{\hbar c}{\sqrt{\alpha} \ell} = \frac{(\mu L - \mu R)c^2}{1 + \frac{8m}{\mu L + \mu R}} \geq 0, \quad \text{gives the regularization bound}$$

$$-\frac{\varepsilon}{2} \leq E \leq \frac{\varepsilon}{2}, \quad \text{where as the squared-mass difference is}$$

$$\Delta \mu_{LR}^2 = (\mu_L^\pm)^2 - (\mu_R^\pm)^2 = 32\pi \frac{\ell p}{\varepsilon} M_p^2.$$

Making use of the right- and left-handed chiral Weyl fields

$$\psi_R = \frac{1 + \gamma^5}{2} \psi, \quad \psi_L = \frac{1 - \gamma^5}{2} \psi, \quad (6)$$

Where solves (2), one gets the system of two massive Weyl equations

$$(\gamma^\mu \hat{p}_\mu + \mu^\pm c^2) \begin{bmatrix} \psi_R^\pm \\ \psi_L^\pm \end{bmatrix} = 0, \quad \mu^\pm = \begin{bmatrix} \mu_R^\pm & 0 \\ 0 & \mu_L^\pm \end{bmatrix} \quad (7)$$

Which can be rewritten as the two-component Schrodinger equation

$$i\hbar \partial_0 \begin{bmatrix} \psi_R^\pm(x, t) \\ \psi_L^\pm(x, t) \end{bmatrix} = -\gamma^0 (i\hbar c \gamma^i \partial_i + \begin{bmatrix} \mu_R^\pm c^2 & 0 \\ 0 & \mu_L^\pm c^2 \end{bmatrix}) \begin{bmatrix} \psi_R^\pm(x, t) \\ \psi_L^\pm(x, t) \end{bmatrix}. \quad (8)$$

For the initial momentum Eigen states  $\psi_{R,L}^\pm(x, t_0)$

$$i\hbar \sigma^i \partial_i \psi_{R,L}^\pm(x, t_0) = p_{R,L}^\pm \psi_{R,L}^\pm(x, t_0), \quad (9)$$

In the Dirac basis, one obtains

$$(\psi_R^\pm)^D(x, t) = \{ [\cos \left[ \frac{t-t_0}{\hbar} E^D(p_R^\pm) \right] - i \mu_\pm^\pm c^2 \frac{\sin \left[ \frac{t-t_0}{\hbar} E^D(p_R^\pm) \right]}{E^D(p_R^\pm)}] \exp \left\{ -\frac{i}{\hbar} p_R^\pm (x-x_0)_i \sigma^i \right\} (\psi_R^\pm)^D(x_0, t_0) - ip_L^\pm c \frac{\sin \left[ \frac{t-t_0}{\hbar} E^D(p_L^\pm) \right]}{E^D(p_L^\pm)} \exp \left\{ -\frac{i}{\hbar} p_L^\pm (x-x_0)_i \sigma^i \right\} (\psi_L^\pm)^D(x_0, t_0) \} \times \exp \left\{ -i \frac{(\mu_R^\pm - \mu_L^\pm)c^2}{2\hbar} (t-t_0) \right\}, \quad (10)$$

And

$$(\psi_L^\pm)^D(x, t) = \{ [\cos \left[ \frac{t-t_0}{\hbar} E^D(p_L^\pm) \right] + i \mu_\pm^\pm c^2 \frac{\sin \left[ \frac{t-t_0}{\hbar} E^D(p_L^\pm) \right]}{E^D(p_L^\pm)}] \exp \left\{ -\frac{i}{\hbar} p_L^\pm (x-x_0)_i \sigma^i \right\} (\psi_L^\pm)^D(x_0, t_0) - ip_R^\pm c \frac{\sin \left[ \frac{t-t_0}{\hbar} E^D(p_R^\pm) \right]}{E^D(p_R^\pm)} \exp \left\{ -\frac{i}{\hbar} p_R^\pm (x-x_0)_i \sigma^i \right\} (\psi_R^\pm)^D(x_0, t_0) \} \times \exp \left\{ -i \frac{(\mu_R^\pm - \mu_L^\pm)c^2}{2\hbar} (t-t_0) \right\}, \quad (11)$$

Where  $E^D(p_R^\pm) = c^2 \sqrt{\left( \frac{\mu_R^\pm + \mu_L^\pm}{2} \right)^2 + \left( \frac{p_R^\pm}{c} \right)^2}$ , while in the Weyl basis

$$(\psi_R^\pm)^W(x, t) = \{ \cos \left[ \frac{t-t_0}{\hbar} E^W(p_R^\pm) \right] - ip_R^\pm c \frac{\sin \left[ \frac{t-t_0}{\hbar} E^W(p_R^\pm) \right]}{E^W(p_R^\pm)} \} \exp \left\{ -\frac{i}{\hbar} p_R^\pm (x-x_0)_i \sigma^i \right\} (\psi_R^\pm)^W(x_0, t_0) + i \mu_L^\pm c^2 \frac{\sin \left[ \frac{t-t_0}{\hbar} E^W(p_L^\pm) \right]}{E^W(p_L^\pm)} \exp \left\{ -\frac{i}{\hbar} p_L^\pm (x-x_0)_i \sigma^i \right\} (\psi_L^\pm)^W(x_0, t_0), \quad (12)$$

And

---

**\*Corresponding author:** Lukasz Andrzej Glinka, Member of the American Association of International Researchers, New York, USA, Tel: +48-660-019-237; E-mail: [laglinka@gmail.com](mailto:laglinka@gmail.com)

**Received** February 11, 2014; **Accepted** February 11, 2014; **Published** February 18, 2014

**Citation:** Glinka LA (2014) On Neutrino in Deformed Relativity. J Astrophys Aerospace Technol 2: e107. doi:[10.4172/2329-6542.1000e107](https://doi.org/10.4172/2329-6542.1000e107)

**Copyright:** © 2014 Glinka LA. This is an open-access article distributed under the terms of the Creative Commons Attribution License, which permits unrestricted use, distribution, and reproduction in any medium, provided the original author and source are credited.

$$(\psi_L^\pm)^W(x, t) = \{ \cos\left[\frac{t-t_0}{\hbar} E^W(p_L^{\pm 0})\right] - \\ + i p_L^{\pm 0} c \frac{\sin\left[\frac{t-t_0}{\hbar} E^W(p_L^{\pm 0})\right]}{E^W(p_L^{\pm 0})} \} \exp\left\{-\frac{i}{\hbar} p_L^{\pm 0} (x-x_0)_i \sigma^i\right\} (\psi_L^\pm)^W(x_0, t_0) + (13) \\ + i \mu_R^\pm c^2 \frac{\sin\left[\frac{t-t_0}{\hbar} E^W(p_R^{\pm 0})\right]}{E^W(p_R^{\pm 0})} \exp\left\{-\frac{i}{\hbar} p_R^{\pm 0} (x-x_0)_i \sigma^i\right\} (\psi_R^\pm)^W(x_0, t_0).$$

Where  $E^W(p_R^{\pm 0}) \equiv c^2 \sqrt{\left(\sqrt{\mu_R^+ \mu_L^+}\right)^2 + \left(\frac{p_R^{\pm 0}}{c}\right)^2}$ .

In this manner, throughout the massive Weyl equation and manifestly unitary two-component Schrodinger equations, the model of Deformed Relativity leads to a massive normalizable neutrino states.

## References

1. Glinka LA (2012) Aethereal Multiverse: A New Unifying Theoretical Approach to Cosmology, Particle Physics, and Quantum Gravity. Great Abington, UK.